Overview of MPLNET Version 3 Cloud Detection

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14 ABSTRACT

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The National Aeronautics and Space Administration Micropulse Lidar Network Version 3 cloud detection algorithm is described and its differences relative to the previous version highlighted. Clouds are identified from normalized Level 1 signal profiles using two complementary methods. The first considers signal derivatives vertically for resolving low-level clouds. The second, which resolves high-level clouds like cirrus, is based on signal uncertainties given the relatively low signal-to-noise ratio exhibited in the upper troposphere by eye-safe network instruments, especially during daytime. Furthermore, a multi-temporal averaging scheme is used to improve cloud detection under conditions of weak signal-to-noise. Diurnal and seasonal cycles of cloud occurrence frequency based on one year of measurements at the Goddard Space Flight Center (Greenbelt, MD) site are compared for the new and previous versions. The largest differences, and perceived improvement, in detection occurs for high clouds (above 5-km, mean sea level) which increase in occurrence by nearly 6%. There is also an increase in the detection of multi-layered cloud profiles from 9% to 20%. Macrophysical properties and estimates of cloud optical depth are presented for a transparent cirrus dataset. However, the limit to which molecular signal can be reliably retrieved above cirrus clouds occurs between cloud optical depths of 0.5 and 0.8.

1. Introduction

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Clouds play a critical role in the Earth's climate system because they are inextricably linked to the hydrological cycle and radiation budget (Liou 1986; Ramanthan et al. 1989). Information about cloud height, thickness, occurrence, and amount are critical inputs for a host of numerical applications involving climate research. Therefore, it is important to have highly accurate and quantitative data records of cloud properties that span several years and geographic regions. Verification of even the most basic modeling processes demands compulsory observations of global cloud occurrence, if there is to be any confidence in their fidelity. Various methods of determining cloud climatologies exist, each with their own advantages and limitations. Visual observations from the surface (Warren et al. 1985; Hahn et al. 1996; Hahn and Warren 1999) provide cloud fraction and morphological cloud types. However, these can be biased by the quality of technician training, underestimation of high clouds, sparse global coverage, and nighttime bias. Passive radiometric sensors aboard satellites, which are the core input of the International Satellite Cloud Climatology Project (ISCCP; Rossow and Schiffer 1991, 1999), offer a true global representation and have the best (unobstructed) potential view of high clouds. However, these can undersample low-level maritime clouds and underrepresent optically-thin cirrus clouds (Holz et al. 2008). Active sensors, like lidar and radar (Platt et al. 1994; Moran et al. 1998; Wang and Sassen 2001), are the primary tools for observing and profiling cloud vertical structure to high accuracy. When flown aboard satellites, like Cloud Aerosol Lidar and Infrared Pathfinder Satellite Observations (CALIPSO; Winker et al. 2007) and CloudSat (Stephens et al. 2002), active sensors also provide global coverage. Even still, the relatively narrow profiling curtain of current active sensors limits observation densities. In the case of CALIPSO and CloudSat, these

missions provide at most two profiles per 24-hour period over most regions, which limits studies of the diurnal impact of clouds on the Earth system. Fundamentally, an array of remote sensing methods is needed in order to investigate the complexity of clouds (Schiffer and Rossow 1983).

2. Micropulse Lidar Network

The National Aeronautics and Space Administration (NASA) Micropulse Lidar Network (MPLNET; Welton et al. 2001, http://mplnet.gsfc.nasa.gov) is a federated network of micropulse lidar (MPL) systems deployed worldwide in support of basic science and the NASA Earth Observing System program (Wielicki et al. 1995). A benefit of MPLNET is the use of a standardized instrument employing a common data processing algorithm with thorough uncertainty characterization, which allows for straightforward comparisons between sites. With sites in polar, mid-latitude, and tropical regions and continuous day/night, high temporal resolution datasets going back as far as 1999, MPLNET datasets represent a valuable archive for improving our understanding of global cloud macrophysical properties on diurnal, season, and decadal scales.

There have been two versions of MPLNET data processing algorithms to date. The first, referred to as Version 1, was released in 2000. Beginning in 2006, the project transitioned to Version 2 (hereafter V2) data products which are currently available. Version 3 (hereafter V3) data processing algorithms are currently in development.

The V2 Level 1 MPLNET data products contain the system diagnostics (e.g. solar background counts, instrument temperature and energy, etc.), normalized relative backscatter (NRB), and NRB uncertainty. The NRB is reported at one-minute temporal resolution up to 30-km, typically at 75-m vertical resolution (some sites operate at 30-m). The NRB is defined as:

$$NRB(z) = C\beta(z)T^{2}(z)$$
 (1)

where C is the instrument calibration constant, β is the backscatter coefficient from both molecules and particles, T^2 is the corresponding total atmospheric two-way transmittance, and z is the altitude. NRB and the calculation of its uncertainty are discussed by Campbell et al. (2002) and Welton and Campbell (2002). Level 1 data are available in real time with no quality assurance. Cloud base and top heights are identified in the V2 Level 1.5 (real time, no quality assurance) and Level 2 (not real time, quality assured) MPLNET data products. The V2 Level 1.5 data products use temperature and pressure profiles from US Standard Atmospheres (COESA 1976) to determine molecular calculations, while National Centers for Environmental Prediction (NCEP)/National Center for Atmospheric Research (NCAR) Reanalysis (Kalnay et al. 1996) is used for V2 Level 2 data.

MPLNET V2 data products have been used to distinguish cloud presence in a number of scientific investigations to date. For example, Campbell and Sassen (2008) use data from the South Pole to document polar stratospheric cloud occurrence over multiple seasons. Shupe et al. (2011) consider MPLNET measurements at Ny-Ålesund, Norway as context for evaluating Artic cloud properties. Others have investigated cirrus contamination of Aerosol Robotic Network (AERONET) aerosol optical depth in Southeast Asia (Chew et al. 2011; Huang et al. 2011) and globally (Huang et al. 2012). Lolli et al. (2013) use collocated 355/527-nm MPLNET observations to estimate the drizzle droplet size from stratocumulus and stratus clouds.

A new V3 cloud detection algorithm has been developed to improve the quality of MPLNET cloud products. The new algorithm uses a combination of signal-processing techniques and a multi-resolution temporal averaging scheme to resolve cloud boundaries. Meteorological profiles provided by the Goddard Earth Observing System – Version 5 (GEOS-5) Atmospheric General Circulation Model (AGCM; Rienecker et al. 2008; Molod et al. 2012) are used for

molecular calculations. Specifically, the Forward Processing for Instrument Teams (FP-IT) GEOS-5 Version 5.9.1 data are utilized (http://gmao.gsfc.nasa.gov/products). The model data are available at 3-hour intervals over 72 pressure levels at 0.625° longitude and 0.5° latitude resolution. Modeled profiles for this study were subsampled from the GEOS-5 grid containing the Goddard Space Flight Center (GSFC) site location, and interpolated to the MPLNET range and time resolutions (75 m, 1 minute).

The goals of this paper are to describe the new algorithm and demonstrate performance. We outline changes relative to V2 cloud detection and describe how the new algorithm is applied to a variety of cloudy scenes. We apply one year of data collected at the GSFC MPLNET site (38.99° N, 76.84° W, 0.05 km above mean sea level; MSL) to compare V2 and V3 results and highlight the impact of our upgraded techniques through differences in macrophysical cloud properties observed from this location.

3. Cloud detection algorithm description

Examples of daytime and nighttime NRB profiles at GSFC are shown in Fig. 1. Both profiles show high-level clouds with base heights near 10 km and top heights near 13 km, MSL. The daytime NRB profile exhibits relatively lower signal-to-noise compared with the nighttime case due to higher solar background, which makes detection of elevated layers an increasingly difficult task.

a. Version 2 cloud detection

Layers are identified in the V2 cloud detection algorithm by a combination of two retrieval methods applied to the Level 1 data products. The first method requires that the first derivative of the lidar signal exceed a minimum threshold in order to detect a layer. The assumption of strong signal gradients makes this well suited for detecting liquid-phase clouds, which are

frequently at lower levels in the NRB profile and correspond with higher signal-to-noise. This is hereafter referred to as the gradient-based cloud detection method (GCDM). The second method is designed for use in cases of low signal-to-noise ratio (SNR) and relies on uncertainties in the lidar signal. This method uses two tunable thresholds and one objective threshold to identify cloud boundaries, and is hereafter referred to as the uncertainty-based cloud detection method (UCDM).

Given the relatively low SNR exhibited by the MPL in the upper troposphere at base one-minute resolution (primarily during daytime), no single procedure is used to detect all cloud types at all times. Thus the merger of these two methods offers the possibility to retrieve the entire cloud vertical structure to the limit of signal attenuation. We describe the basis for each method, as follows.

1) GRADIENT-BASED CLOUD DETECTION

Autonomous methods of cloud detection using gradients in the lidar signal, such as the differential zero-crossing method described by Pal et al. (1992), are well established. The first step in the GCDM is to normalize the NRB using the attenuated molecular backscatter coefficient,

$$\beta_m'(z) = \beta_m T_m^2(z),\tag{2}$$

which produces an attenuated scattering ratio, β_r , multiplied by the instrument calibration constant as

$$C\beta_{r}'(z) = \frac{C\beta(z)T^{2}(z)}{\beta_{m}'(z)} = \frac{C[\beta_{m}(z) + \beta_{p}(z)]T_{m}^{2}(z)T_{p}^{2}(z)}{\beta_{m}(z)T_{m}^{2}(z)} = C\left[1 + \frac{\beta_{p}(z)}{\beta_{m}(z)}\right]T_{p}^{2}(z).$$
(3)

Here the subscripts m and p denote contributions from molecules and particles, respectively.

The first derivative of $C\beta_r'$ is used to identify clouds in the GCDM. Due to increasing uncertainty in the profile with height, the GCDM retrieval is only performed up to a "noise

altitude", defined as the altitude at which the uncertainty, δ NRB, exceeds half of the NRB (or conversely, analogous to an SNR of 2). Cloud presence corresponds with an increase in total backscatter, which results in a large positive gradient in $C\beta_r$ with height as seen in Fig. 2. The threshold used to identify the cloud base is defined as

$$a_{\text{max}} = K \cdot \overline{C\beta_r'}, \qquad (4)$$

where $\overline{C\beta_r}'$ is the mean value of $C\beta_r'$ up to the noise altitude and K is an empirical parameter (unitless), set to 10 for this study. The value of K is chosen carefully so that it is high enough to reject insignificant peaks in the first derivative of $C\beta_r'$ (i.e. aerosol stratification in the surface-attached layer or signal noise) while remaining sensitive enough to identify weakly-scattering clouds.

The cloud base is identified at the altitude bin immediately preceding that where the first derivative of $C\beta_r$ exceeds a_{max} . Identification of cloud top is more ambiguous and is performed using one of two processes. A negative gradient in $C\beta_r$ occurs near the top of a cloud, corresponding with the decrease in total backscatter and the impact of signal attenuation through the cloud. The first method used to identify the cloud top relies on a threshold defined as

$$a_{\min} = \overline{C\beta_r'} - a_{\max} . \tag{5}$$

The algorithm begins by looking for altitude bins above the cloud base where the first derivative of $C\beta_r'$ falls below a_{\min} . Then the altitude bin where the first derivative initially returns above a_{\min} is identified as the cloud top. However, if this condition is not met, the cloud top is chosen as the altitude bin where the value of $C\beta_r'$ falls below the value at the cloud base or the noise altitude, whichever occurs first. If the lidar signal becomes significantly attenuated within the

cloud, the designation as an apparent cloud top is more appropriate in accordance with standards agreed upon by the Experimental Cloud Lidar Pilot Study (ECLIPS; Platt et al. 1994).

Two cloud layers are apparent in Fig. 2. Both cloud bases are identified at the altitudes immediately below the a_{max} exceedances (dashed line, positive derivative). The first (lowest) cloud top can be found using the a_{min} threshold (dashed line, negative derivative). But the derivative never falls below a_{min} for the second cloud layer. Therefore, the alternative process is used to identify the apparent cloud top. Only true (not apparent) cloud tops are reported in V2 MPLNET cloud products.

2) UNCERTAINTY-BASED CLOUD DETECTION

An alternative to algorithms that utilize gradients in the lidar return to identify clouds are approaches that compare cloudy lidar returns to clear sky returns (Clothiaux et al. 2007). Similarly, the UCDM uses a theoretical molecular return and the signal uncertainty to detect elevated clouds, and is fully described by Campbell et al. (2008, hereafter C08). However, a few comments regarding its implementation are warranted. The first step in the UCDM is to approximate the value of the instrument calibration constant. Level 1 NRB is divided by the attenuated molecular backscatter coefficient, resulting in an attenuated scattering ratio multiplied by the instrument calibration constant, as given in Eq. (3). Next, a clear-sky search is performed to locate a normalization region where we can approximate that β_p approaches zero over a certain number of range bins, N. The nature of the UCDM only allows for cloud detection at altitudes above the normalization region. The calibration constant is approximated by averaging Eq. (3) over the N bins. C08 stress that this final normalization value, C_f^* , must be distinguished from C due to unknown transmission losses below the normalization region.

As an example, a representation of the UCDM is shown in Fig. 3. Beginning at 1-km above ground level (AGL) and working upward, the uncertainty in $C\beta_r'$ is evaluated at each altitude bin according to the criteria set forth by C08 until a "clear-air" slot is found. In this case, the "clear-air" region used to determine the normalization value begins at $r_1 = 6.46$ km and ends at $r_N = 8.71$ km. Averaging $C\beta_r'$ between r_1 and r_N gives a value of $C_f^* = 92.33$ MHz km³ sr μJ^{-1} .

Once C_f^* has been calculated, a so-called pseudo-attenuated backscatter and its uncertainty are solved as

$$PAB(z) = \frac{NRB(z)}{C_f^*}$$
(6)

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$$\delta PAB(z) = PAB(z) \sqrt{\left[\frac{\delta NRB(z)}{NRB(z)}\right]^2 + \left(\frac{\delta C_f^*}{C_f^*}\right)^2} . \tag{7}$$

Eq. (7) is then modified by substituting the attenuated molecular backscatter for the PAB to develop an objective threshold,

$$\alpha(z) = \beta_m(z)T_m^2(z) + \beta_m(z)T_m^2(z)\sqrt{\left[\frac{\delta NRB(z)}{\beta_m(z)T_m^2(z)C_f^*}\right]^2 + \left(\frac{\delta C_f^*}{C_f^*}\right)^2},$$
(8)

- which is used to differentiate what are first presumed cloud returns from molecular return.
- Range bins above r_N that meet the condition

$$PAB(z) - \delta PAB(z) > \alpha(z) \tag{9}$$

are then evaluated to determine whether they represent particulate layer base heights. Range bins that do not meet the condition from Eq. (9) are used to establish "clear-air" slots and are disqualified from the particulate base height search. In other words, a minimum detectable scattering ratio for particulates is defined as

$$\beta_{r,\min}(z) = \frac{\alpha(z) + \delta PAB(z)}{\beta'_{m}(z)}.$$
 (10)

Running averages of PAB and δ PAB are used in conjunction with two additional tunable thresholds, ϕ and κ (both analogous to a SNR), to determine the actual layer base and top heights. The threshold ϕ sets the minimum average value of PAB/ δ PAB for bins that exceed Eq. (10) in order to identify the layer base. At this point, and as described above, the UCDM assumes that any such layers detected are hydrometeor clouds, thereby leaving the potential for false detection of elevated aerosol layers. In the absence of supplementary information, however, such as color ratio (Liu et al. 2005) or depolarization (Cho et al. 2008, Omar et al. 2009) and combined with the goal of resolving as much thin cirrus as possible in the low SNR portions of the NRB profile, this is unavoidable. Mitigation strategies are described further below. The threshold κ sets the minimum average value of PAB/ δ PAB for bins that do not exceed Eq. (10) in order to identify clear air layers and consequently particulate layer tops. In Fig. 3, the red line indicates the threshold used to distinguish particulate from molecular returns. Bins that exceed this objective threshold are evaluated using the tunable threshold ϕ , while bins with values lower than this threshold are evaluated using the tunable threshold κ . The sensitivity of V2 cloud detection to the tunable thresholds is evaluated in C08 and the values chosen for ϕ and

3) V2 CLOUD RETRIEVALS

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 κ will depend on the site location and instrument performance parameters.

An example of V2 cloud retrievals is shown in the top panel of Fig. 4. Results from the GCDM and UCDM are integrated based on the noise altitude (described in Section 3.a.1). Clouds occurring below this height are reported from the GCDM. All clouds above the noise altitude are identified using the UCDM. At night, the noise altitude reaches above typical cirrus cloud heights at GSFC. Therefore, the GCDM is almost exclusively responsible for cloud detection. As a result, weakly-scattering cirrus can go undetected, since GCDM thresholds are tuned primarily with boundary layer phenomena in mind (i.e. suppression of aerosol identification). This can be observed frequently between 0300 and 0600 UTC in Fig. 4, where cirrus presence is underreported and cloud base heights are overestimated.

In the daytime, the noise altitude shown in Fig. 4 falls between 8–9 km and the UCDM is responsible for all cloud detection above it. In several instances, cloud bases (red markers) are shown while the corresponding cloud tops (orange markers) appear to be missing. In these cases the lidar signal is assumed to be significantly attenuated, and therefore no cloud top is reported.

b. Version 3 cloud detection

The V3 algorithm is based on V2 with a few meaningful changes to the UCDM. Consequently, the changes in V3 represent an update to C08 and how the GCDM and UCDM are merged. A schematic of the V3 cloud detection algorithm at the one-minute base NRB temporal resolution is shown in Fig. 5. Low-altitude obstructions (e.g. fog or low stratus decks) reduce SNR and limit the accuracy of cloud retrievals. Therefore, each profile is screened for these "beam-blocked" conditions by a process described further below. If no such obstructions are found, the first step in the UCDM is to calculate the normalization value.

1) NORMALIZATION REGION

As mentioned in C08, it is most practical to find a normalization region to calculate C_f^* nearest to the instrument in order to increase the depth of the profile analyzed for clouds. In V2,

the clear-sky search is initiated at 1 km AGL. However, in V3, an altitude of 5-km MSL is chosen to reduce the likelihood of normalizing within relatively homogenous aerosol layers in or near the boundary layer. Furthermore, based on recent global cloud and aerosol studies (Holz et al. 2008; Sassen et al. 2008; Campbell et al. 2015), 5 km represents a reasonable height to begin the search for high clouds in tropical and mid-latitude locations. A lower altitude may be needed for polar sites, however. In the event that it is not possible to perform the normalization step at 5-km, the UCDM attempts to normalize lower in the atmospheric profile iteratively, to as low as 1 km AGL, as done in the V2 algorithm.

The normalization region also serves as the boundary between the GCDM and UCDM retrievals in the V3 cloud algorithm, allowing the better-suited method (GCDM for low clouds and UCDM for high clouds) to operate during both day and night. The V3 cloud retrievals in the bottom panel of Fig. 4 can be compared with the V2 retrieval in the same figure to see the relative apparent improvement. We also note that there are conditions when either method (GCDM or UCDM) may be used to retrieve high-level or low-level clouds. Therefore, the retrieval method for each cloud layer is provided as an output parameter.

2) OBJECTIVE THRESHOLD

The objective UCDM threshold, α , defined in Eq. (8) is now adjusted for attenuation within cloud layers in V3. In order to attenuate the α threshold, we assume an initial unity transmission at the base of the first detected cloud layer. Then the two-way transmittance is calculated at each altitude bin Z within the cloud layer as

$$T_{C}^{2}(Z) = T_{C}^{2}(Z-1)\exp\left[-2S_{C}\beta_{C}(Z)\Delta z\right]$$

$$T_{C}^{2}(Z) = T_{C}^{2}(Z-1)\exp\left[-2S_{C}\left(\frac{\beta_{r}'(Z)}{T_{C}^{2}(Z-1)}-1\right)\beta_{m}(Z)\Delta z\right],$$
(11)

where β_C is the cloud backscatter coefficient, S_C is the extinction-to-backscatter ratio (or lidar ratio) of the cloud, Δz is the vertical resolution of the instrument, and β_r' is the attenuated scattering ratio given by

$$\beta_r'(z) = \frac{\text{NRB}(z)}{C_f^* \beta_m'(z)} \approx \left[1 + \frac{\beta_C(z)}{\beta_m(z)} \right] T_C^2(z).$$
 (12)

Overestimating the single unknown term in these equations, the extinction-to-backscatter ratio, will lead to excessive attenuation of the objective threshold and eventual mischaracterization of clear sky as cloud. Therefore, in this step, a very conservative choice for S_C equal to the molecular extinction-to-backscatter ratio ($8\pi/3$) is used.

Once the transmittance has been determined, Eq. (8) is modified to recalculate the α threshold as

$$\alpha(z) = \beta_m(z) T_m^2(z) T_C^2(z) + \beta_m(z) T_m^2(z) \sqrt{\left[\frac{\delta NRB(z)}{\beta_m(z) T_m^2(z) C_f^*}\right]^2 + \left(\frac{\delta C_f^*}{C_f^*}\right)^2} .$$
 (13)

- By attenuating the α threshold, we are able to better resolve cloud tops for optically-thick clouds.
- Furthermore, multilayer clouds, where the upper layer was missed entirely by the V2 algorithm,
- are now more effectively detected in some cases.
- 275 3) MULTI-TEMPORAL RESOLUTION
- Cloud boundaries are only reported at one-minute temporal resolution in the V2 algorithm.
- However, instances of high solar background reduce UCDM performance. So, as described by
- 278 C08, multi-temporal resolution settings are used in V3. In addition to the one-minute base
- temporal resolution, the UCDM is performed for intermediate (five-minute) and long (twenty-
- 280 minute) temporal averages using a sliding window centered on a corresponding one-minute
- profile.

A flow chart describing the process is shown in Fig. 6. Retrievals at the base temporal resolution are used to screen profiles at longer averages, thus limiting the effects of "beamblocked" profile contamination. Within a window of N profiles, k profiles are removed from the average if an attenuating structure is detected below 5 km at the base temporal resolution. If k > N/2, then the entire average profile is rejected.

A combined cloud scene is created using cloud boundaries retrieved from the three temporal resolutions. First, the combined scene uses the cloud boundaries from the base temporal resolution. Next, the intermediate and then long temporal resolutions are used to fill in the missing gaps as needed. Cloud boundaries are always reported at the highest temporal resolution possible to reduce the effects of cloud smearing caused by averaging. This is broadly consistent with the methodology used for NASA Cloud Aerosol Lidar with Orthogonal Polarization (CALIOP) Level 2 products and their gridding of multiple spatial resolutions from the selective iterated boundary locator (SIBYL) to their vertical feature mask (Vaughan et al. 2005). A noted difference is that SIBYL also uses an intensity-clearing process to remove features detected at finer resolutions from the coarser spatial averages. No such intensity clearing is performed with the V3 algorithm.

4) FALSE POSITIVES

As mentioned previously, use of the UCDM presumes first only cloud presence. Additional constraints are thus used to reduce the number of instances when noise excursions, elevated aerosol layers or poor normalizations produce false cloud retrievals. The first constraint establishes a minimum layer thickness of 150 m (i.e., two bins at 75-m resolution) in order to qualify a potential layer as a cloud. Therefore, we require all bins within a distance greater or equal to the minimum layer thickness to exceed the minimum detectable scattering ratio in Eq. (10) before a cloud base is established. Similarly, a minimum clear air distance of 150 m is used

to avoid falsely segmenting a single cloud into multiple layers. All bins within the minimum clear distance must fall below the minimum detectable scattering ratio in order to establish a cloud top. The second constraint requires that the standard deviation of the attenuated scattering ratio (β_r) within the detected layer exceed an empirically-determined threshold, σ_{\min} , which varies as a function of cloud top temperature. Cloud layers generally cause large variances in β_r , either through attenuation effects in otherwise colloidally-stable liquid water clouds or since ice crystals fall within cirrus cloud layers, creating complex structures. By contrast, aerosols in the free troposphere settle in stratified stable layers absent of convection and are expected to be homogeneous within each layer. Thus, clouds layers can be distinguished by their relatively large standard deviations of β_r .

In order to determine σ_{min} , a dataset was developed consisting of 144 days at GSFC in 2012 and 27061 retrieved UCDM layers at one-minute resolution (18308 thin cirrus cloud layers, 3233 non-cirrus layers, 5520 aerosol layers) when the particulate type could be reasonably identified from visual inspection (Fig. 7). Thin cirrus clouds are distinguished using a cloud top temperature threshold of -37 °C (Sassen and Campbell 2001, Campbell et al. 2015) and a maximum cloud optical depth (COD) of 0.3 (Sassen and Cho 1992). The COD calculation uses a process described by Chew et al. (2011) and is discussed fully in Section 3.c.1. Non-cirrus clouds are those with cloud top temperature warmer than -37 °C.

Several choices for σ_{min} were evaluated using error matrices (Congalton and Meade 1986) and the corresponding values of accuracy and Matthews correlation coefficient (Matthews 1975; MCC) which are defined as

$$Accuracy = \frac{TP + TN}{TP + FP + FN + TN} \tag{14}$$

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$$MCC = \frac{TP \times TN - FP \times FN}{\sqrt{(TP + FP)(TP + FN)(TN + FP)(TN + FN)}},$$
(15)

where TP is the number of instances when clouds were correctly identified, TN is the number of instances when aerosols were misidentified as clouds, and FN is the number of instances when clouds were misidentified as aerosols. The choice for σ_{min} that resulted in the highest values of accuracy (0.92) and MCC (0.74) is given by

$$\sigma_{\min} = \begin{cases} 2, & \text{for } T > -37 \,^{\circ}\text{C} \\ 10^{\frac{T+40}{10}}, & \text{for } -47 \,^{\circ}\text{C} < T < -37 \,^{\circ}\text{C} \\ 0.2, & \text{for } T < -47 \,^{\circ}\text{C} \end{cases}$$
 (16)

The error matrix for the empirically-determined σ_{min} is provided in Table 1. However, we note that the dataset contains two atypical long-range smoke transport events (with corresponding top height temperatures between -50 °C and -60 °C), which comprised 35% of the aerosol category. If these two events are removed, the accuracy and MCC improve to 0.98 and 0.92, respectively. Similar results were observed at the five- and twenty-minute resolutions, but not shown for brevity.

The final constraint used to distinguish cloud from aerosol layers is that the estimated COD exceed a threshold, τ_{min} . Through empirical testing, we estimate COD and set $\tau_{min} = 0.005$ based on analysis of these subsets relative to the perception of how noise impacts these sub-samples combined with a similar analysis by Thorsen et al. (2011).

We briefly note here that lidars with polarization capabilities have recently been incorporated into the MPLNET project. However, because the overwhelming majority of

existing data (which spans from 2000 – present) was collected without polarization, the algorithm presented here does not rely on such data. It remains as a future goal to demonstrate how polarization can be used to improve aerosol-cloud discrimination, once a sufficient amount of data is collected from the new polarized sites.

c. Version 3 algorithm output

A listing of the V3 cloud detection algorithm output parameters are provided in Table 2. The output parameters from all temporal averages are gridded to one-minute temporal resolution, as previously described in the combined cloud scene. The number of cloud layers detected, day flag, and attenuation altitude are given as a single value each minute, characterizing the atmospheric column. All other cloud products and data flags correspond with individual cloud layers, and are provided each minute with dimensions equal to the number of cloud layers detected. Meteorological values at the cloud boundaries are obtained from the interpolated GEOS-5 profiles described in Section 2.

1) CLOUD PHASE AND CIRRUS CLOUD OPTICAL DEPTH

In the absence of visual cloud observations, as is the case for autonomous lidar measurements made by MPLNET, Sassen and Campbell (2001) recommend using a minimum cloud top temperature of -37 °C to identify cirrus. In the V3 cloud algorithm, we use this thermal threshold to distinguish ice clouds (i.e. cirrus) from all other cloud phases. Because depolarization capabilities are not standard for all MPLNET instruments, no attempt is made to distinguish liquid from mixed phase clouds.

Campbell et al. (2015) evaluate the -37 °C cloud top temperature threshold globally versus the Level 2 CALIOP algorithms that identify ice-phase cloud layers and found that over 99% of clouds satisfying this thermal threshold were classified as ice. Furthermore, 81% of all ice clouds had cloud top temperatures less than -37 °C. They conclude, consistent with the findings

of Sassen and Campbell (2001), that this thermal threshold is stable for specifically distinguishing cirrus cloud presence in lidar studies that lack depolarization, though there is some ambiguity in cases of "warm" cirrus that likely coincide with cloud top temperature greater than -37 °C.

An estimated COD is calculated for clouds distinguished as cirrus using the procedure described by Chew et al. (2011). Two-way cloud transmission is calculated using Eq. (11). However, now the value of S_C is selected based on the cloud top temperature. Reported values of S_C are on the order of 16–18 sr for liquid water clouds (Pinnick et al. 1983, Yorks et al. 2011) and 10–40 sr for cirrus (Sassen and Comstock 2001; Chen et al. 2002; Yorks et al. 2011; Garnier et al. 2015). A value of $S_C = 18$ sr is chosen for layers with clouds top temperatures warmer than -37 °C and $S_C = 20$ sr at colder temperatures where cirrus clouds are expected. We note that due to uncertainty in the lidar ratio for cirrus clouds, these estimates may represent the lower limit of COD.

Next, Eq. (12) is used to solve for the cloud backscatter coefficient and the estimated COD is given by

$$\tau = S_C \int_{base}^{top} \beta_C(z) dz = S_C \int_{base}^{top} \beta_m(z) \left(\frac{\beta_r'(z)}{T_C^2(z)} - 1 \right) dz.$$
 (17)

The fidelity of the COD estimate is limited by the choice of S_C and accuracy of cloud boundaries retrieved. The relative error in the lidar-derived optical depth is smallest for low optical depths and proportional to $\Delta S_C/S_C$ as τ approaches zero (Winker et al. 2009). Lidar signals are unable to penetrate through optically-thick clouds, which causes uncertainty in the value at the apparent cloud top. In these cases, the estimated COD will be biased low. Similarly, attenuation from the bottom-most cloud layer leads to uncertainty in corresponding retrievals of higher clouds for

lidar profiles containing multiple cloud layers. As a result, the estimated COD will be most reliable for single-layer, optically-thin clouds.

2) RETRIEVAL INDEX

While cloud boundaries are only reported at a single temporal resolution, a retrieval index is included to indicate whether the cloud was also detected at one or more of the other temporal averages. Cloud layers at different temporal resolutions are considered the same if (i) they share a common base or top height within a vertical depth of 250 m or (ii) one cloud layer is completely enveloped within the other.

An example of a combined cloud scene, with corresponding retrieval indices, is shown in Fig. 8. The value of the retrieval index is equal to the sum of the temporal resolutions used to identify the cloud layer. For example, if a cloud is detected at all three temporal resolutions, the value of the retrieval index is 1+5+20=26. The advantage of the multi-temporal averaging scheme can be seen during the day between 1400–1500 UTC in Fig. 8. The elevated cloud layer (~15 km) is mostly undetected at the one-minute resolution, but can be resolved using the longer averages. The cloud layer at ~2 km produces "beam-blocked" conditions that prevent use of higher-temporal averages for much of the cirrus cloud layer above it.

3) ATTENUATION ALTITUDE

Because the lidar signal can become completely attenuated within optically-thick clouds, it is important to determine when a true cloud top is being reported as opposed to an apparent cloud top. Nadir-pointing lidar instruments have an advantage of using the ground return to determine if the lidar signal has been extinguished. However, with zenith-pointing lidar, that determination is more tenuous. Winker and Vaughan (1994) defined a transmittance index to determine when the lidar signal was fully attenuated based on the percentage of samples above the cloud top that exceeded the background. Other techniques used for zenith-pointing lidar

have included the use of a minimum threshold lidar signal along with its slope (Wang and Sassen 2001) and comparisons with molecular profiles (Lo et al. 2006).

In V3, cloud tops (both true and apparent) are reported for all cloud layers along with the altitude at which the lidar signal is determined to be fully attenuated. This attenuation altitude is found by starting at the range bin of the highest reported cloud altitude and incrementally moving upwards in the profile until, within a depth of 2 km, (i) the percent difference between the mean pseudo-attenuated backscatter and modeled attenuated molecular backscatter falls below some threshold, T1, and (ii) either the backscatter signal falls below a minimum value or the percentage of range bins where the backscatter signal is less than zero exceeds a threshold T2.

This application pertains specifically to profiles that contain clouds or other obstructions, since the attenuation thresholds can also be satisfied by other conditions that lead to low SNR (e.g. high solar background). "Beam-block" conditions from low-altitude obstructions are found

with the same search criteria, though the search is limited to the first 2 km above the surface.

4. Results

In order to demonstrate the effects of the changes implemented in the V3 algorithm, we compare V2 and V3 cloud retrievals for one year at the GSFC MPLNET site. Table 3 and Fig. 9 show data sampling statistics for 2012, including the total number of profiles and percentage of time when 1-minute NRB measurements were available monthly. Observable profiles are given as the number and percentage of available profiles that are not "beam-blocked" below 2 km, MSL. Profile attenuation was determined using the V2 method because it is the most restrictive and ensures an even comparison between the two cloud detection algorithms. The diurnal distribution of data recorded and successful V3 normalizations to calculate C_f^* are also shown. There is very little differentiation between the distributions of V2 and V3 normalizations, so only

the number of successful V3 normalizations are shown for simplicity. Though MPLNET measurements are continuous, there is a decrease in the number of successful normalizations near solar noon in relatively low SNR conditions.

Because V3 uses a merged cloud scene and V2 is only processed at one-minute resolution, V3 retrievals are evaluated using the base one-minute resolution (hereafter V3b) and the merged cloud scene (hereafter V3m). Comparisons are limited to cloud base statistics because cloud tops are not recorded for all V2 retrievals. Finally, we describe the macrophysical and optical characteristics of cirrus clouds observed during this study, again adhering to the methodology described in Campbell et al. (2015), using V3 retrievals.

a. Vertical dependence

Figure 10 shows the cloud base distributions retrieved from the V2 and V3 algorithms, respectively, at GSFC during 2012. A bimodal distribution similar to that observed by Winker and Vaughan (1994), with peaks at ~1-2 km and ~9-10 km is apparent. The total numbers of cloud observations are 269505, 304363, and 332810 for the V2, V3b, and V3m retrievals, respectively. Compared with V2, the number of cloud observations increases by 12.9% and 23.5% for V3b and V3m, respectively. The largest increase in the number of clouds observed occurs at altitudes above 5 km.

Because the difference in the number of clouds retrieved shows a clear vertical dependence, we examine them specifically for three sub-samples, by defining low clouds as those with base heights less than 2 km, high clouds as those with base heights greater than 5 km, and middle clouds as those with base heights between 2-5 km (WMO 1975). The number of lidar profiles for each classification, along with occurrence frequency, is shown in Table 4. Cloud occurrence frequency is defined as the number of lidar profiles containing a particular cloud classification

divided by the total number of observable profiles. Regardless of the retrieval method (V2, V3b, and V3m), occurrence frequency is nearly identical for low clouds, which reflects the relative consistency in GCDM application between V2 and V3 at one-minute resolution. High clouds show the largest increase in occurrence frequency. For example, comparing the V2 and V3m algorithms, the occurrence frequency of high clouds increases by 5.9% attributable to (i) the increased identification of elevated, multi-layer cloud decks using an attenuated UCDM threshold, (ii) increased use of the UCDM to identify high clouds at day and night, and (iii) multi-temporal application of UCDM to increase SNR. To (i), V2 retrievals resulted in 91% of cloudy lidar profiles containing single-layer clouds. The percentage of single-layer clouds decreases to 83% and 80% for V3b and V3m, respectively.

471 b. Seasonal dependence

Figure 11 shows the annual cycle for low, middle, high, and total cloud classifications during 2012. The low cloud occurrence frequency is nearly identical for all three retrieval methods. Middle clouds retrieved using V3b and V3m exhibit a slight separation from V2. The largest differences are again seen with high-cloud retrievals. While the annual cycles for high clouds show similar patterns for all three retrievals, there is an increase in occurrence frequency of ~3% and 6% for V3b and V3m, respectively. The increase in high-cloud occurrence frequency when compared to V2 ranges from 1% to 4% using V3b and 4% to 10% for V3m. The largest differences for high-cloud occurrence frequency between V2 and V3 occurs during summer months, which is coincident with the period when the sun is at its highest elevation and thus solar background is highest.

c. Diurnal dependence

Differences in the diurnal cycle show similar characteristics as the annual cycles for low and middle clouds. As seen in Fig. 12, V2, V3b, and V3m are nearly identical for low clouds. While V3b and V3m show slight differences from V2, they are indistinguishable from each other. High-cloud diurnal cycles follow the same trends for all three retrievals. However, the cloud occurrence frequency is higher for V3b and highest for V3m. No clear diurnal trend is apparent at GSFC because some changes (e.g. using the UCDM at all times) affect both day and night retrievals. At tropical sites, where the solar background is higher and longer temporal averaging is necessary, there may be a more obvious diurnal trend.

d. Macrophysical and optical cirrus properties

Based on the greater detection of high clouds demonstrated above, we characterize cirrus clouds over the GSFC site as detected by the V3 algorithm. As stated earlier, cirrus presence is determined using a cloud top temperature threshold of -37 °C. Additionally, we limit the analysis to cases when (i) only cirrus clouds (no underlying liquid water or mixed phase clouds) were detected in the profile, (ii) the estimated COD was less than 3, based on the upper-limit for cirrus clouds suggested by Sassen and Cho (1992), and (iii) the attenuation altitude was at least 2 km above the cloud top. The final constraint limits the analysis to "transparent cirrus" cases for which the algorithm is more likely to identify the true cloud top.

The resulting dataset includes 57930 cirrus clouds. The majority of cloud detections (82%) occur at the base one-minute temporal resolution. The largest occurrence rate of the coarse temporal averages occurs at or near noon and during the summer months when the solar background is highest.

Table 5 summarizes the seasonal and annual mean characteristics of the transparent cirrus dataset. The monthly variation in the macrophysical properties is shown in Fig. 13. Cirrus

clouds over GSFC tend to be higher and thinner (geometrically and optically) in the spring and summer and lower and thicker in the fall and winter seasons. Cirrus also occur more frequently in the spring and summer months. The transparent cirrus dataset is composed almost entirely (~95%) of sub-visual (COD < 0.03) and thin (COD < 0.3) cirrus clouds. Uncertainties in the value of the extinction-to-backscatter ratio and cloud top height could lead to an exaggeration of this finding. However, it should be noted that when the extinction-to-backscatter ratio is increased from 20 sr to 30 sr (not shown), 86% of cirrus clouds still have a COD less than 0.3. These results are qualitatively consistent with the findings of Dupont et al. (2010), who reported 50-75% of non-opaque cirrus clouds had an optical thickness less than 0.3 based on ground-based lidar and CALIPSO observations at four mid-latitude sites.

Frequency distributions of the optical and macrophysical properties are presented in Fig. 14. The cloud optical depth peaks in the sub-visual range and has a positive skew. The transparent cirrus dataset suggests that the limit at which we are able to resolve molecular signal above cloud, and thus reliably determine the cloud top, occurs near a COD of 0.5. However, if an extinction-to-backscatter ratio of 30 sr is used, this limit occurs near a COD of 0.8.

A comparison of daytime and nighttime cloud retrievals is provided in Table 6. There are only slight differences in the occurrence frequency between day and night cases. However, the geometric and optical depths are considerably lower in the daytime. The thinning of daytime cirrus may be attributable to difficulty in correctly identifying cloud boundaries due to solar background effects (Thorsen et al. 2013). However, convective cloud remnants are also likelier to occur during daytime hours. Decoupling the two, aside from seasonal influence, is outside the scope of this analysis. In the same manner, the daytime retrievals are more likely to be

considered as totally attenuated due to the higher solar background reducing the possibility to resolve molecular signal at cirrus heights.

5. Summary and Discussion

A new Version 3 (V3) cloud detection algorithm has been developed within the NASA Micropulse Lidar Network (MPLNET) that uses a combination of retrieval methods and a multitemporal averaging scheme. Most V3 changes represent updates to the Version 2 (V2) uncertainty-based threshold algorithm introduced by Campbell et al. (2008). The threshold used to identify cloud presence now accounts for attenuation losses within cloud layers, which allows for better estimation of cloud tops and boundaries of overlying cloud layers in profiles where multiple cloud layers are detected. A more synergistic merging of the gradient-based cloud detection method (GCDM) and uncertainty-based cloud detection method (UCDM) improves nighttime clouds detection of tenuous high clouds. The incorporation of coarser temporal resolutions at intermediate (5-minute) and long (20-minute) averages improves detection in situations with low SNR (e.g. high solar background). One year of data at the NASA Goddard Space Flight Center (GSFC) in Greenbelt, MD is used to show the effect of these updates on cloud retrievals.

The largest impact of the changes to the cloud detection algorithm is evident with high clouds (those with cloud base > 5 km), while the diurnal and annual cycles of low and middle clouds exhibit only slight changes from V2 to V3. The high-cloud occurrence frequency increases by nearly 6% at GSFC when using the V3 merged cloud scene compared with the V2 retrieval. Furthermore, the ability to detect multi-layered cloud scenes is improved with the V3 algorithm. The results show that 91% of clouds in 2012 at the NASA GSFC project site were recorded as single-layer clouds according to the V2 retrieval compared with 80% for V3.

A brief investigation of macrophysical properties for transparent cirrus clouds shows that the mean base and top heights at GSFC occur at 10.17 ± 1.63 km and 11.07 ± 1.43 km, respectively. The highest and thinnest (both geometrically and optically) cirrus are found during the spring and summer months, which was coincident with the highest cirrus occurrence frequency. There is no significant difference in occurrence frequency between daytime and nighttime retrievals. However, cirrus clouds are thinner (both geometrically and optically) in daytime than nighttime, which may be attributed to a combination of increased uncertainty due to the solar background effects and higher occurrence of convective cloud remnants during the day. Notably, the limit to which we are able to resolve molecular signal above cirrus clouds occurs between cloud optical depths of 0.5 and 0.8, allowing for uncertainty in the extinction-to-backscatter ratio.

The value of the MPLNET cloud datasets is in its continuous (both day and night) and long-term measurements at polar, mid-latitude, and tropical sites using a standard instrument and data processing algorithm. Incorporating the V3 cloud retrievals from MPLNET as part of a multi-instrument investigation will enhance our current knowledge of clouds, in particular cirrus. As it stands, the cloud products provide a unique validation dataset for the modeling community and satellite measurements. With some MPLNET sites now well into their second decade of continuous cloud and aerosol observations, the project has become an integral component of ground-based evaluation of atmospheric processes and verification of NASA satellite missions. This paper thus represents our continuing effort to optimize the fidelity of project datasets for the benefit of the community and in sustaining general scientific inquiry.

Acknowledgments.

The authors acknowledge Larry Belcher for processing the V2 lidar data, and the MPLNET PIs and staff for their efforts in establishing and maintaining the GSFC site. The GEOS-5 meteorological data were provided by the NASA Global Modeling and Assimilation Office (GMAO) at GSFC. The NASA Micro-Pulse Lidar Network is funded by the NASA Earth Observing System and Radiation Sciences Program. Author JRC acknowledges the support of NASA Interagency Agreement NNG13HH10I on behalf of MPLNET.

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TABLE 1. Error matrix for cloud-aerosol discrimination using σ_{min}

	Predicted cloud	Predicted aerosol
True cloud	21242	299
True aerosol	1851	3669

734 Table 2. MPLNET V3 cloud detection algorithm output

Parameter	Description	
Base and top altitudes	Cloud boundaries in km above MSL	
N _{layers}	Number of cloud layers detected	
N _{profiles}	Number of 1-min profiles used in average profile	
Method of cloud detection	GCDM or UCDM	
Cloud phase	Ice or liquid/mixed phase based on -37 °C threshold	
Estimated cirrus COD	Calculated using lidar ratios of 20 and 30 sr, ranging from	
	0.005 to 3	
Base and top temperatures	From FP-IT GEOS-5 Version 5.9.1	
Base and top pressures	From FP-IT GEOS-5 Version 5.9.1	
Base and top wind speeds	From FP-IT GEOS-5 Version 5.9.1	
Base and top wind directions	From FP-IT GEOS-5 Version 5.9.1	
Day/night flag	0 = Night, 1=Day (from ephemeris)	
Retrieval index	Indicates temporal average used to identify the cloud layer	
Attenuation altitude	Altitude where lidar beam is deemed to be significantly	
	attenuated	

TABLE 3. Summary of data collected at GSFC in 2012 (percentages shown in parentheses)

zations	V3 normali-	profiles	Observable	recorded	Profiles	
(78)	34798	(85)	38022	(100)	44615	Jan
(84)	34882	(92)	38226	(100)	41690	Feb
(79)	35248	(84)	37303	(100)	44611	Mar
(84)	36090	(93)	39971	(100)	43142 32095	Apr
(74)	23644	(81)	25870	(72)		Apr May Jun
(86)	35801	(92)	38138	(96)	41479	Jun
(84)	34303	(87)	35339	(91)	40817	Jul
(89)	39408	(88)	39318	(100)	44485	Aug
(86)	37158	(88)	38104	(100)	43078	Sep
(71) (84) (70)	20649	(81)	38104 23412	100) (65) (99) (99)	29081	Oct
(84)	36224	(96) (79)	41265	(99)	42946	Nov
	36224 31041 399246	(79)	41265 35175 468165		43078 29081 42946 44345 536999	Sep Oct Nov Dec Year
(74)	399246	(87)	468165	(93)	536999	Year

736 TABLE 4. Number of lidar profiles and occurrence frequency at GSFC in 2012

V2	V3b	V3m	
	Low clouds		
77943	78511	78513	
0.181	0.183	0.183	
	Middle clouds		
69612	75353	75462	
0.162	0.175	0.175	
	High clouds		
109366	120750	134514	
0.254	0.281	0.313	
	Total clouds		
246132	255079	267777	
0.572	0.593	593 0.623	

TABLE 5. Transparent cirrus cloud properties

	Spring	Summer	Fall	Winter	Annual
	(MAM)	(JJA)	(SON)	(DJF)	
Cloud base					
Height (km)	10.46 ± 1.44	11.17 ± 1.48	9.77 ± 1.33	9.13 ± 1.49	10.17 ± 1.63
Temperature (°C)	-50.3 ± 9.7	-47.9 ± 9.2	-43.9 ± 8.9	-45.7 ± 10.2	-47.4 ± 9.9
Pressure (mb)	262.6 ± 57.6	243.9 ± 52.0	291.0 ± 53.7	312.4 ± 66.8	275.9 ± 63.7
Wind speed (m s ⁻¹)	23.3 ± 12.5	20.3 ± 11.0	26.6 ± 10.7	38.6 ± 14.5	27.1 ± 14.2
Wind direction (°)	277.2 ± 40.2	236.8 ± 110.3	259.4 ± 28.3	279.1 ± 20.5	264.6 ± 63.3
Cloud top					
Height (km)	11.39 ± 1.35	11.93 ± 1.34	10.74 ± 1.16	10.09 ± 1.33	11.07 ± 1.48
Temperature (°C)	-57.0 ± 8.1	-53.5 ± 8.0	-51.3 ± 7.3	-52.9 ± 8.4	-54.1 ± 8.3
Pressure (mb)	225.9 ± 47.0	215.4 ± 41.6	249.9 ± 41.2	267.6 ± 52.4	238.4 ± 50.5
Wind speed (m s ⁻¹)	24.5 ± 12.9	21.8 ± 11.8	29.6 ± 11.4	42.5 ± 15.6	29.3 ± 15.4
Wind direction (°)	277.2 ± 37.4	238.3 ± 108.0	259.5 ± 29.6	279.3 ± 20.6	265.0 ± 61.7
Cloud depth (km)	0.93 ± 0.61	0.76 ± 0.51	0.97 ± 0.70	0.96 ± 0.69	0.90 ± 0.63
Estimated COD	0.07 ± 0.10	0.06 ± 0.10	0.09 ± 0.14	0.09 ± 0.12	0.08 ± 0.11
Cirrus Type (%)					
Sub-visual	47	51	45	41	46
Thin	50	46	49	53	49
Opaque	3	3	6	6	5
Occurrence (%)	17	12	10	13	13
Cloud layers	18664	13889	10769	14608	57930

739 TABLE 6. Daytime and nighttime transparent cirrus properties

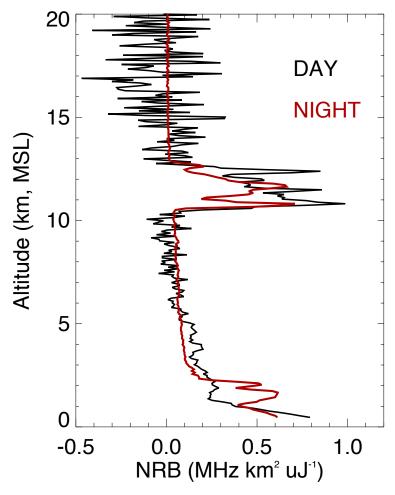
	Daytime	Nighttime
Cloud base		
Height (km)	10.15 ± 1.69	10.19 ± 1.58
Temperature (°C)	-47.3 ± 10.0	-47.5 ± 9.8
Pressure (mb)	277.2 ± 65.8	274.9 ± 61.9
Wind speed (m s ⁻¹)	27.1 ± 14.0	27.0 ± 14.5
Wind direction (°)	258.6 ± 70.4	269.6 ± 56.2
Cloud top		
Height (km)	10.94 ± 1.51	11.19 ± 1.44
Temperature (°C)	-53.2 ± 8.4	-54.8 ± 8.1
Pressure (mb)	243.5 ± 51.4	234.0 ± 49.3
Wind speed (m s ⁻¹)	29.1 ± 15.1	29.5 ± 15.6
Wind direction (°)	259.8 ± 68.3	269.4 ± 55.2
Cloud depth (km)	0.79 ± 0.56	1.00 ± 0.67
Estimated COD	0.07 ± 0.11	0.09 ± 0.12
Cirrus Type (%)		
Sub-visual	51	42
Thin	46	52
Opaque	3	6
Occurrence (%)	12	14
Cloud layers	26360	31570

740 **List of Figures**

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- boundaries encloses 50% of the data nearest the median and thick boundaries enclose 90% of the data. The dashed line indicates the empirically-determined σ_{\min} .
- 8 (Top) Example of a combined cloud seen at the GSFC MPLNET site on 7 July 52 2012. (Bottom) Retrieval flags showing the temporal resolutions used to detect the combined cloud scene. Square symbols are used to indicate the cloud base and the Xs are used for the cloud top.
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- Annual cycle of the occurrence frequency for low, middle, high and total clouds at GSFC during 2012. The solid line represents V2, the dashed line V3b, and the dash-dotted line V3m.
- Diurnal cycle of the occurrence frequency for low, middle, high and total clouds at GSFC during 2012. The solid line represents V2, the dashed line V3b, and the dash-dotted line V3m.
- Monthly averaged cloud top (thick solid line), cloud base (thin solid line), cloud 57 depth (dashed red line) and cloud optical depth (dashed-dotted blue line) for the transparent cirrus dataset.
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is 0.5 km for base and top altitudes and 0.25 km for cloud depth. (d) Frequency distribution (black) and cumulative frequency distribution (red). The vertical dashed lines indicate the thresholds for sub-visual (0.03) and thin (0.3) cirrus clouds. Horizontal axis bin size equals 0.001.



743 FIG. 1. Examples of the NRB at GSFC on 7 July 2012 at day (1800 UTC, black) and night (0300 UTC, red).

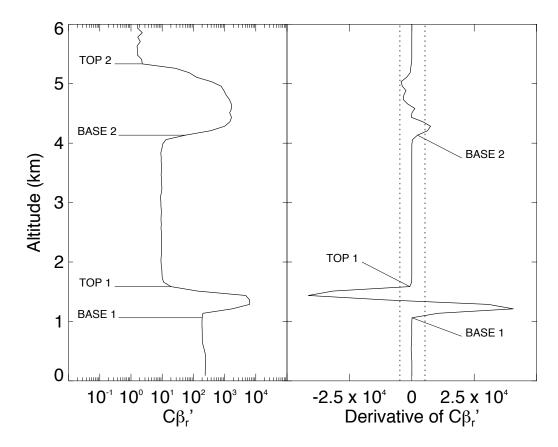


FIG. 2. Representation of the GCDM (GSFC; 0233 UTC 7 Nov 2012). The product of the calibration constant and attenuated scattering ratio is shown in the left panel and its derivative is shown on the right. The vertical dashed lines represent the a_{\min} and a_{\max} thresholds.

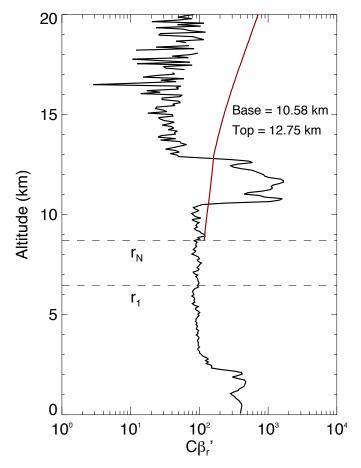


FIG. 3. Representation of the UCDM (GSFC; 0300 UTC 7 July 2012). The bottom (r_1) and top (r_N) of the normalization region are indicated by the horizontal dashed lines. The red curve represents the product of C_f^* and the minimum detectable scattering ratio used to determine the cloud boundaries.

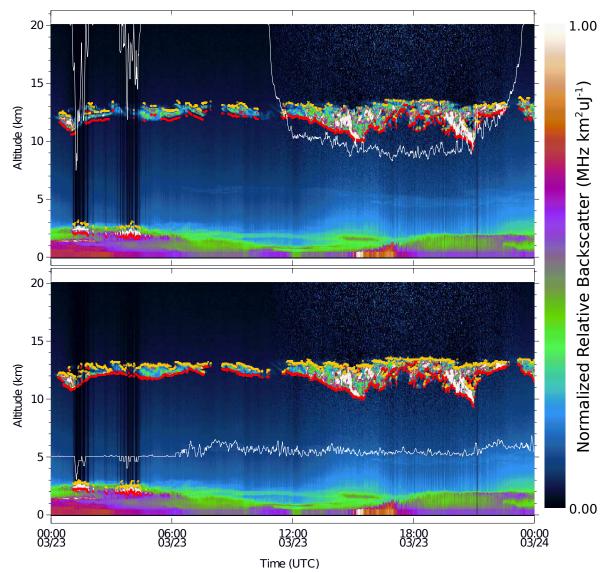
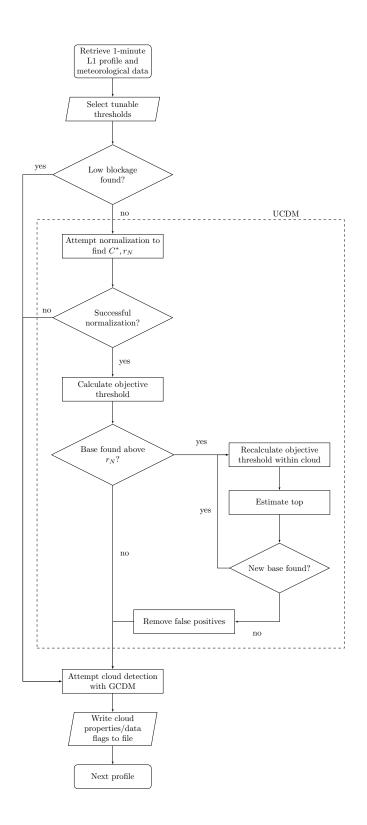


FIG. 4. (Top) Version 2 cloud retrievals at GSFC site on 23 March 2012 (one-minute resolution). Cloud bases and tops are indicated by red and orange markers, respectively. The noise altitude (i.e. δ NRB/NRB > 0.5) is represented by the solid white line (four-minute smoothing applied for clarity). (Bottom) Same as top figure but using the V3 algorithm at one-minute resolution. The bottom of the normalization region is represented by the solid white line (four-minute smoothing applied for clarity).



761
 762 FIG. 5. Schematic of the V3 cloud detection algorithm at the one-minute base resolution.

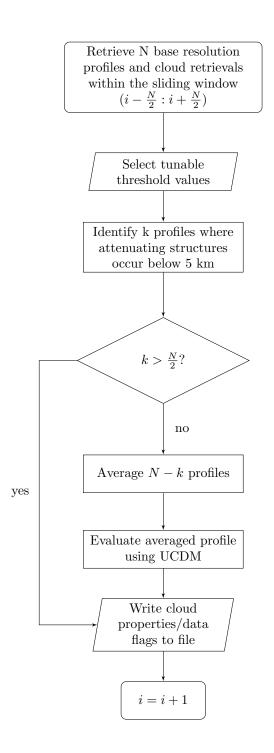


FIG. 6. Schematic of the V3 cloud detection algorithm for higher temporal averages.

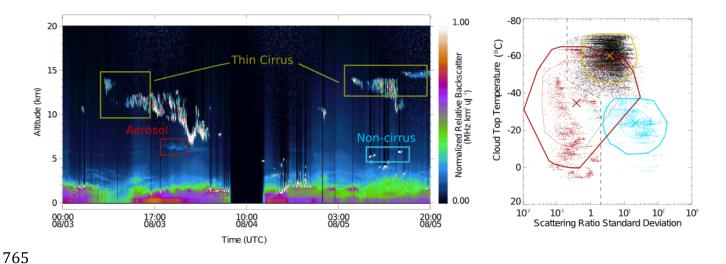


Fig. 7. (Left) Representative groupings of cirrus clouds, non-cirrus clouds, and aerosols from GSFC (3-5 Aug 2012). (Right) Scatterplot of full dataset used to determine σ_{min} threshold. The Xs represent the median values of each group, the thin boundaries encloses 50% of the data nearest the median and thick boundaries enclose 90% of the data. The dashed line indicates the empirically-determined σ_{min} .

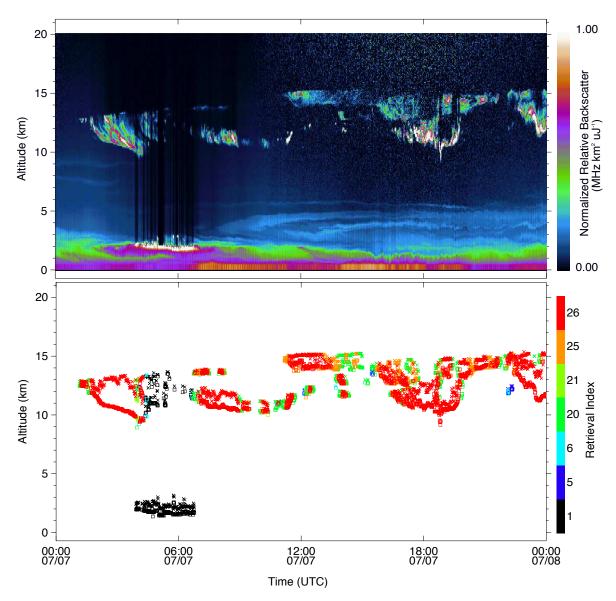


FIG. 8. (Top) NRB from lidar measurements at the GSFC MPLNET site on 7 July 2012. (Bottom) Retrieval indices showing the temporal resolutions used to detect the combined cloud scene. Square symbols are used to indicate the cloud base and the Xs are used for the cloud top.

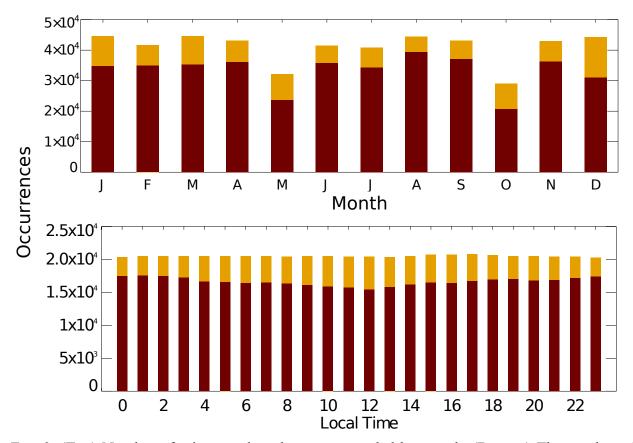


FIG. 9. (Top) Number of minutes when data was recorded by month. (Bottom) The number of minutes data was recorded each hour. The red colors are the successful C_f^* normalizations and yellow are the total data.

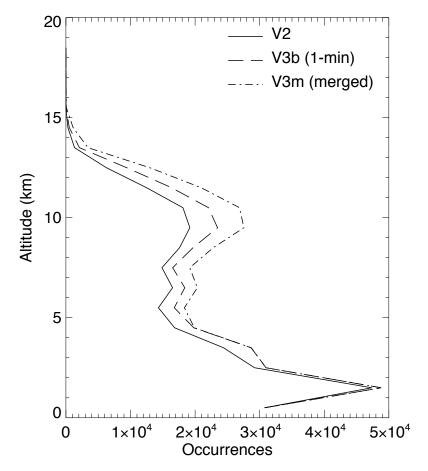


FIG. 10. Cloud base height distributions for GSFC during 2012 for V2 (solid line), V3b (dashed line) and V3m (dash-dotted line). Vertical axis bin size equals 1 km.

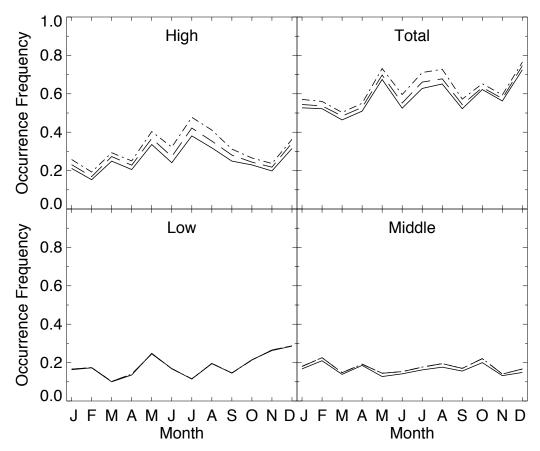


FIG. 11. Annual cycle of the occurrence frequency for low, middle, high and total clouds at GSFC during 2012. The solid line represents V2, the dashed line V3b, and the dash-dotted line V3m.

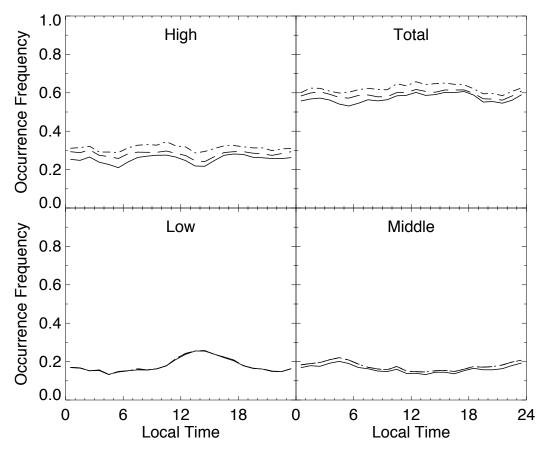


FIG. 12. Diurnal cycle of the occurrence frequency for low, middle, high and total clouds at GSFC during 2012. The solid line represents V2, the dashed line V3b, and the dash-dotted line V3m.

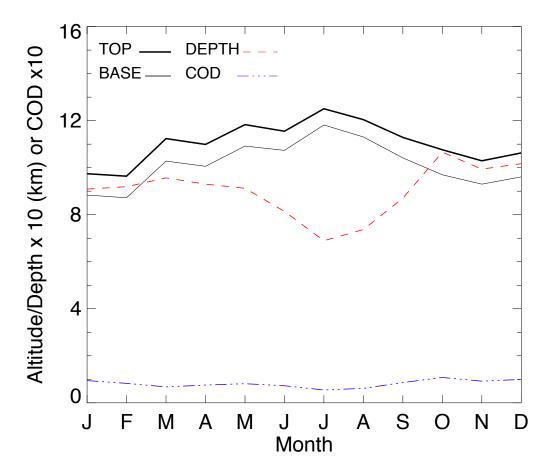


Fig. 13. Monthly averaged cloud top (thick solid line), cloud base (thin solid line), cloud depth (dashed red line) and cloud optical depth (dashed-dotted blue line) for the transparent cirrus dataset.

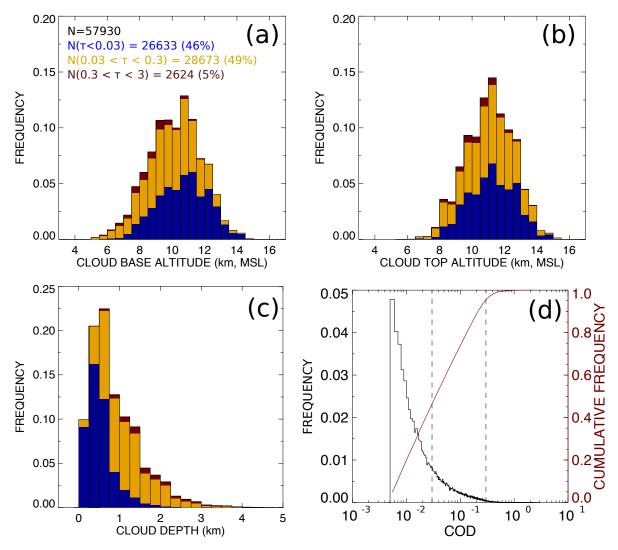


FIG. 14. (a) Cloud base altitude, (b) cloud top altitude and (c) cloud depth for the transparent cirrus dataset. The colors indicate the cirrus type based on the estimated COD (blue for subvisual, yellow for thin, and red for opaque cirrus). Horizontal axis bin size is 0.5 km for base and top altitudes and 0.25 km for cloud depth. (d) Frequency distribution (black) and cumulative frequency distribution (red). The vertical dashed lines indicate the thresholds for sub-visual (0.03) and thin (0.3) cirrus clouds. Horizontal axis bin size equals 0.001.